

A new Lorentzian Geometry

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Fakultät für Mathematik



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FWF Österreichischer
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excellent = austria

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From smooth Riemannian to synthetic Lorentzian
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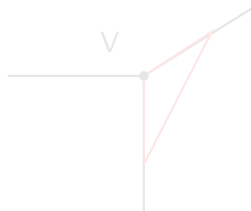
The larger picture: Regularity in geometry

Regularity determines basic geometric features

Example 1: Walking on a sphere vs. walking on a cube.



Sphere: Walking along a geodesic/meridian is locally always the shortest path.

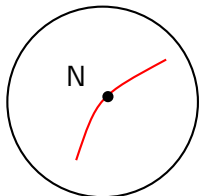


Cube: It is always shorter to deviate to the right face than to go along the edge.

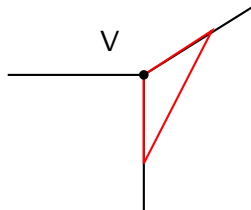
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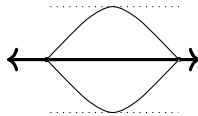
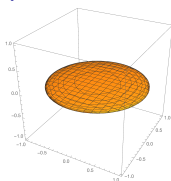
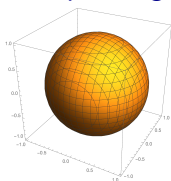
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Example 2: Squeezing the sphere.

[Hartman-Wintner, 52]



Convexity fails for metrics of Hölder regularity $g \in C^{1,\alpha}$ ($\alpha < 1$):

Equator still geodesic but it is always shorter to deviate into hemispheres

Realised by $M := (-1, 1) \times \mathbb{R}$ with

$$g_{(x,y)} = (1 - |y|^\lambda)dx^2 + dy^2 \quad (1 < \lambda < 2)$$

Lorentzian version: [Sämman-S, 19]

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Lorentzian Geometry is the Language of General Relativity

GR, Einstein's theory of space, time & gravity:

Gravity is universal, hence prop. of surrounding spacetime

Curvature proportional to **mass/energy content**



$$\text{Ric} - \frac{1}{2} R g = \frac{8\pi G}{c^4} T \quad (\text{E})$$

Geometric description via a spacetime manifold (M, g)

- $M \dots$ (4-dimensional) smooth manifold
- $g \dots$ *smooth Lorentzian metric* on M :
non-degenerate scalar product g_p in each tangent space $T_p M$
with signature $(-, +, +, +)$ **not positive definite!**

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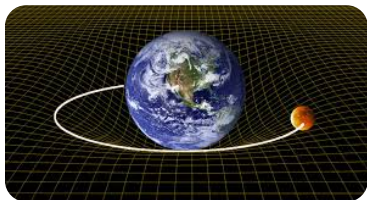
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spacetime tells matter how to move.

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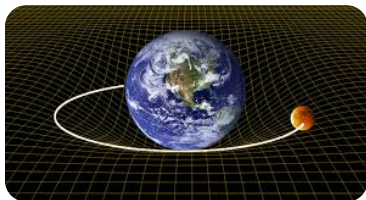
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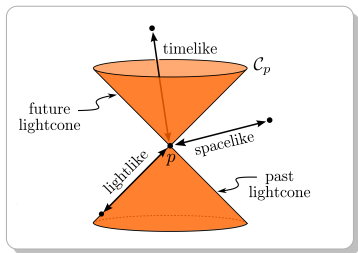
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Causality is the fundamental structure of Lor. Geometry



$$v \in T_p M, \quad g_p(v, v) \quad \left\{ \begin{array}{ll} < 0 & \text{timelike} \\ = 0 & \text{lightlike/null} \\ \leq 0 & \text{causal} \\ > 0 & \text{spacelike} \end{array} \right.$$

extends naturally to
(loc. Lipschitz) curves

Causal relations between points $p, q \in M$

$p \ll q : \Leftrightarrow$ connected by future timelike curve

$p \leq q : \Leftrightarrow$ connected by future causal curve

Chronological/causal future/past

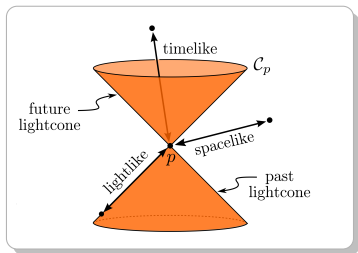
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past versions: $I^-(p), J^-(p)$

for $A \subseteq M$: $I^\pm(A), J^\pm(A)$

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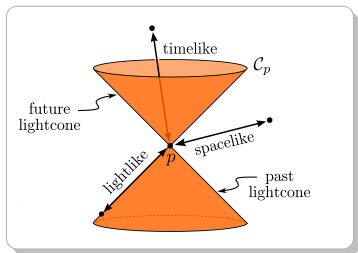
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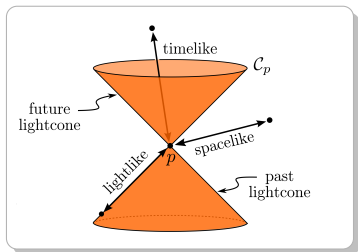
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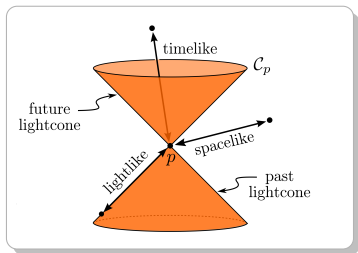
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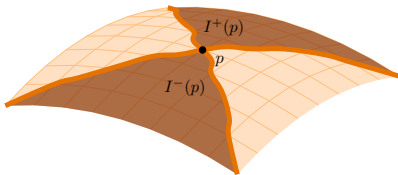


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Smooth Lorentzian Geometry is not enough

Generally from (E): $Ric - \frac{1}{2}Rg = \frac{8\pi G}{c^4} T$; $Ric \propto \partial^2 g + (\partial g)^2$;
 $T \notin C^0 \rightsquigarrow g \propto C^{1,1}$; below C^2 (smooth f.a.p.p.) **non-smooth:=below C^2**

Motivations from physics & mathematics

- *PDE/initial value* point-of-view (regularity in hyperbolic PDE!)
- physically relevant *models*: matched spacetimes
[Senovilla, Mars, Sánchez-Pérez, Manzano, Ohanian, S,...]
impulsive gravitational waves
[Barrabés, Griffiths, Hogan, Podolský, Sämann, Švarc, S....]
- approaches to *Quantum Gravity* (causal set theory, CDT,...)

"The discontinuities of the matter [...] which must be allowed [...] pose enormous problems [...]. To avoid this annoying problem though—despite it being completely fundamental!—we will implicitly assume that g is at least C^2 ."
[Garcia Parrado-Senovilla, CQG, 2005]

But much has happened since!

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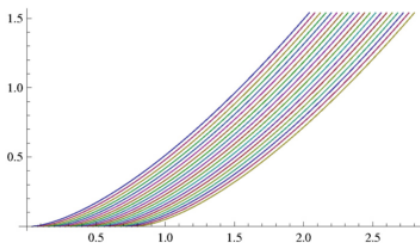
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Back to larger picture: Regularity in Lorentzian causality

Regularity determines basic geometric features

Example 3: Lightcones bubble up.

[Chruściel-Grant, 12]



Some null geodesics for $\alpha = 1/3$

$$M = \mathbb{R}^2, g \in C^{0,\alpha} \ (\alpha < 1)$$

$$g = -(dt + (1 - |t|^\alpha)dx)^2 + dx^2$$

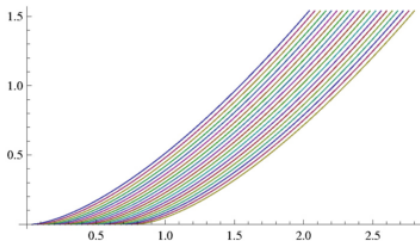
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 $\leadsto$ null cone has full measure
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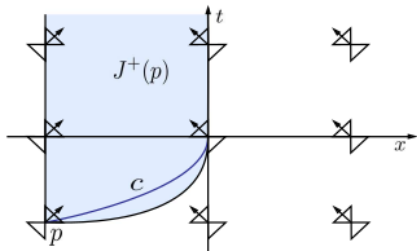
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Ex. 4: The future is not always open. [Grant-Kunzinger-Sämann-S, 20]



$$M = \mathbb{R}^2, g \in C^{0,\alpha} \ (\alpha < 1)$$

The blue curve c is timelike
but reaches $\partial I^+(p)$

$\leadsto I^+(p)$ is *not open*

Lightcones turn in non-Lipschitz way

What can be done: Approaches to non-smooth LG

Spacetime setting: smooth manifold but metric g below C^2 (*analytic*)

- distributional curvature: $g \in H^1 \cap L^\infty$ maximal w stable $\mathcal{D}'$ -curvature
[Geroch-Traschen, 87], [LeFloch-Mardare, 07], [Vickers-S, 09]
- non-linear distributional geometry (algebras of generalised functions)
[Farkas, Grosse, Hanel, Hörmann, Kunzinger, Mayerhofer, Nitsch,
Oberhuber, Spreitzer, Vickers, S. . . 01–]
- advanced regularisation techniques $\check{g}_\varepsilon \prec g \prec \hat{g}_\varepsilon$ [Chruściel-Grant, 12]
- causality for $g \in C^0$ [Sämman, 16], closed cone structures [Minguzzi, 19]
- ↪ Singularity theorems in low regularity: $g \in C^{1,1}$, $g \in C^1$, $g \in C^{0,1}$
[Graf, Grant, Kunzinger, Ohayan, Schinnerl, Stojkovic, Vickers, S, 15–]
currently: $g \in W^{1,p}$, $R \in L^p$ [Kunzinger-Reintjes-Vega-S, 25–]
- C^0 -inextendibility [Sbierski, 15–, Ling, 24, Monsani-Solanki-Prados-Sämman, ong.]

A new geometry: beyond the manifold setting (*synthetic*)

This will be the focus for the rest of the talk.

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How to detect curvature: A glimpse of Riemannian mflds

Riemannian manifold (M, h) ; curvature $R_{XY}Z = [\nabla_X, \nabla_Y]Z - \nabla_{[X,Y]}Z$

$$\text{Sectional curvature } \text{Sec}(X, Y) = \frac{\langle R_{XY}Y, X \rangle}{\|X\|^2\|Y\|^2 - \langle X, Y \rangle^2}$$

Theorem (Toponogov $\triangle$ -comparison) $\text{Sec} \iff$

For all (small) geodesic triangles $\triangle abc$ in (M, h) consider a comparison triangle $\triangle \bar{a}\bar{b}\bar{c}$ in the .

Then for all p, q on its sides and corresponding comparison points $\bar{p}, \bar{q}$

$$d_h(p, q) \leq d(\bar{p}, \bar{q}).$$

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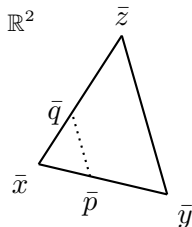
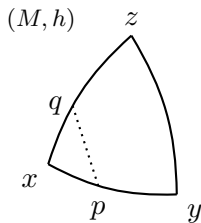
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For all (small) **geodesic triangles** $\triangle abc$ in (M, h) consider a **comparison triangle** $\triangle \bar{a}\bar{b}\bar{c}$ in the Euclidean Plane $\mathbb{R}^2$.

Then for all p, q on its sides and corresponding **comparison points** $\bar{p}, \bar{q}$

$$d_h(p, q) \geq d_{\mathbb{R}^2}(\bar{p}, \bar{q}).$$



Riemannian distance d_h

$$d_h(p, q) := \inf \{ L_h(\gamma) : \gamma \text{ from } p \text{ to } q \}$$

w length of locally Lipschitz curves γ

$$L_h(\gamma) = \int \sqrt{h(\dot{\gamma}, \dot{\gamma})}$$

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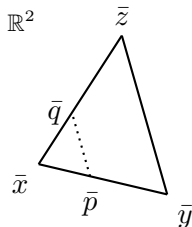
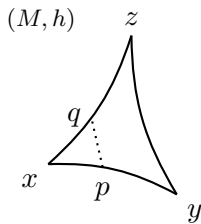
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Riemannian manifold (M, h) ; curvature $R_{XY}Z = [\nabla_X, \nabla_Y]Z - \nabla_{[X,Y]}Z$

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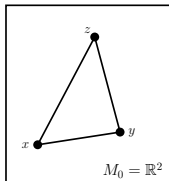
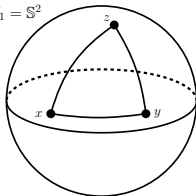
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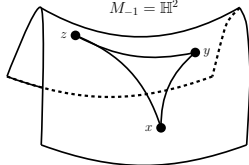
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model spaces M_K of constant curvature K

$M_1 = \mathbb{S}^2$



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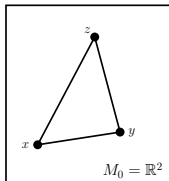
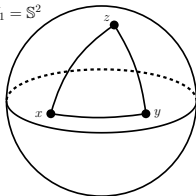
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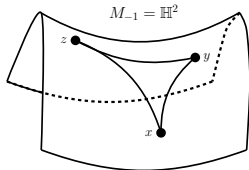
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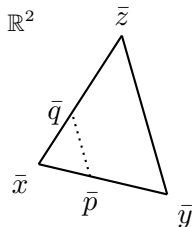
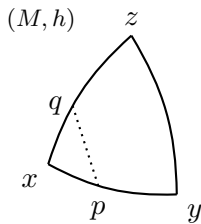
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Triangle condition

- needs no manifold structure, only distances between pts.
- works on metric spaces, can be turned into definition

Sectional curvature bounds for metric spaces

Definition (Length space) (X, d) metric space w d intrinsic, i.e.

$$d(x, y) = \inf\{L_d(\gamma) : \gamma \text{ from } x \text{ to } y \text{ continuous}\}$$

Geodesics $\gamma : [0, 1] \rightarrow X$ with $L_d(\gamma) = d(\gamma(0), \gamma(1))$

Definition (Synthetic curvature bounds)

A length space has curvature bounded by K if for all (small) *geodesic triangles* $\triangle abc$ and their *comparison triangles* $\triangle \bar{a}\bar{b}\bar{c}$ in the model space of curvature K and all points p, q on its sides and corresponding $\bar{p}, \bar{q}$

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Curvature bounded below / above: *Alexandrov spaces* / *CAT(K)-spaces*

Metric geometry: rich theory since the 1980-ies: GH-convergence, preservation of curvature bds., Gromov pre-compactness thm.

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Definition (“Correct” curvature bounds) [Andersson-Howard, 98]

A smooth Lorentzian manifold has $\text{Sec} \geq K$ if

spacelike sectional curvatures $\geq K$ and *timelike* sectional curvatures $\leq K$.

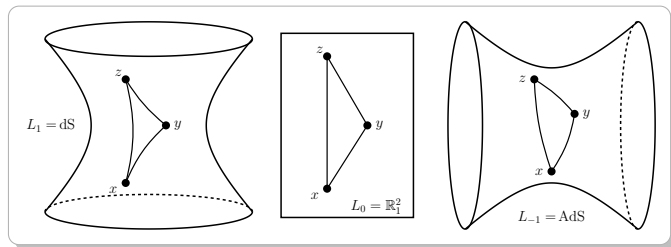
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*Lorentzian
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- $K > 0$: de Sitter
- $K = 0$: Minkowski
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Go synthetic!

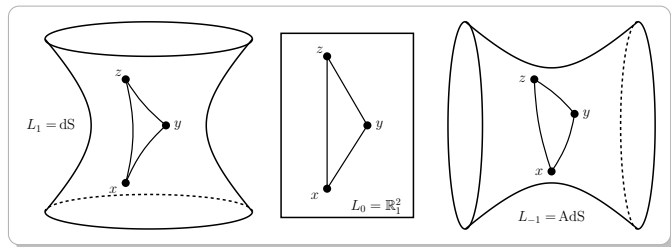
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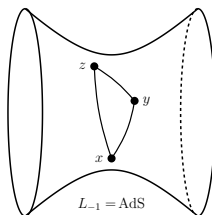
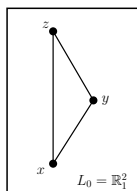
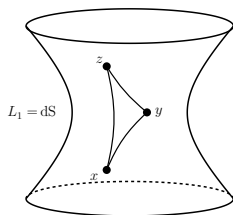
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How to go beyond Lorentzian manifolds?

		smooth		synthetic
pos. def.		Riemannian mfls		metric length spaces
Lorentzian		Lorentzian manifolds		???

What is the analogue of metric (length) spaces in the *Lorentzian setting*?

Serious issue

No natural distance function $\leadsto$ no direct entry into world of metric spaces

Lorentzian (pre-)length spaces [Kunzinger-Sämman, 18] based on causal spaces [Kronheimer-Penrose, 67] & the use of the **time separation**

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Lorentzian time separation vs. Riemannian distance

Length of causal curves $\gamma : I \rightarrow M$: $L_g(\gamma) := \int_I \sqrt{-g(\dot{\gamma}, \dot{\gamma})}$

Null curves have $L_g(\gamma) = 0$, timelike can be arbitrarily short!

Time separation function

$$\tau(p, q) := \sup\{L_g(\gamma) : \gamma \text{ future causal from } p \text{ to } q\}$$

Properties:

- lower semi-continuous,
- satisfies reverse Δ -inequality:

$$\tau(p, q) + \tau(q, r) \leq \tau(p, r)$$

- τ is **not** a distance function

Why: implements basic causality already of SRT (twin paradox)

Consequences: No cheap entry into the world of metric geometry!

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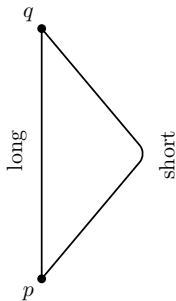
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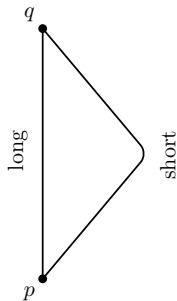
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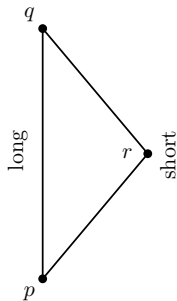
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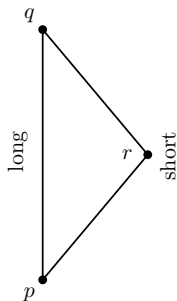
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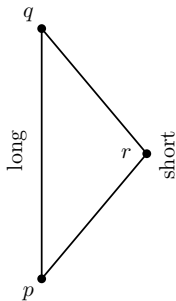
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Causality w/o metric: Lorentzian (pre-)length spaces

Causal space: X (metrizable) topological space with **abstract causality:**
 $\leq$ preorder on X , $\ll$ transitive relation contained in $\leq$
Abstract time separation: $\tau: X \times X \rightarrow [0, \infty]$ lower semicontinuous

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It is a **Lorentzian length space** if τ is intrinsic.

Examples

- **smooth spacetimes** (M, g) with usual time separation τ
- **Lorentz(-Finsler) spacetimes**, of **low regularity** ($g \in C^0 + \dots$)
- **finite directed graphs** (causal sets)

Lorentzian **causality theory**

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It is a *Lorentzian length space* if τ is intrinsic.

Examples

- *smooth spacetimes* (M, g) with usual time separation τ
- *Lorentz(-Finsler)* spacetimes, of *low regularity* ($g \in C^0 + \dots$)
- *finite directed graphs* (causal sets)

Lorentzian *causality theory*

variants by [Braun-McCann, 23-], [Minguzzi-Suhr, 24-], [Müller, 24-]

Causality w/o metric: Lorentzian (pre-)length spaces

Causal space: X (metrizable) topological space with *abstract causality:*
 $\leq$ preorder on X , $\ll$ transitive relation contained in $\leq$

Abstract time separation: $\tau: X \times X \rightarrow [0, \infty]$ lower semicontinuous

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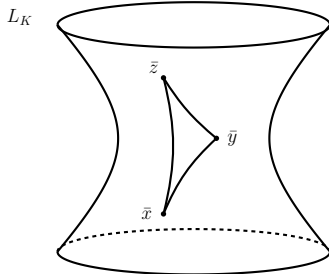
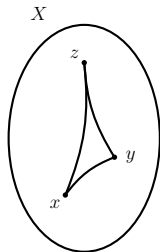
Timelike curvature via triangle comparison

Definition (Synthetic curvature bounds)

$(X, \ll, \leq, \tau)$ has *timelike curvature* $\geq K$ if

- 1 some technical conditions hold
- 2 for all *small timelike triangles* Δxyz and their comparison $\Delta \bar{x}\bar{y}\bar{z}$ in L_K and all p, q resp. $\bar{p}, \bar{q}$

$$\tau(p, q) \leq \bar{\tau}_K(\bar{p}, \bar{q}).$$



Faithful extension of
sectional curvature bounds
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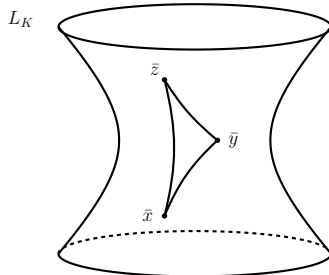
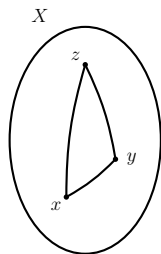
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Timelike curvature comparison: The whole truth

Theorem (Equivalence of curvature conditions)

[Beran-Kunzinger-Rott, 24]

Let X be a chronological LpLS. Recall curvature bds. below/above by K in the following senses:

- | | |
|---|--|
| i Timelike triangle comp. | ix Angle comp. |
| ii One-sided timelike triangle comp. | x Hinge comp. |
| iii Causal triangle comp. | xi Timelike four point condition |
| iv One-sided causal triangle comp. | xii Angle version of timelike four point condition |
| v Strict causal triangle comp. | xiii Causal four point condition |
| vi One-sided strict causal triangle comp. | xiv Strict causal four-point condition |
| vii Monotonicity comp. | xv τ -convexity (resp. τ -concavity) condition |
| viii One-sided monotonicity comp. | |

If X is strongly causal, regular, and locally D_K -geodesic, and in the case of upper curvature bounds additionally is locally causally closed and in the case of lower curvature bounds satisfies (8), then *all notions of curvature bounds are equivalent*.

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In general the following relations hold:

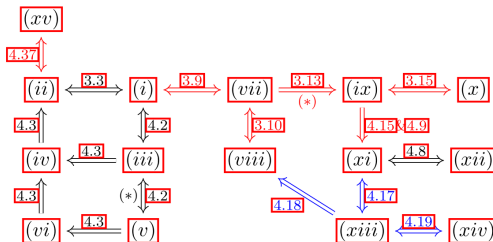


Figure 4: All relations between different formulations of curvature bounds for Lorentzian pre-length space. Black arrows are always valid, red arrows require X to be regular, and blue arrows require X to be strongly causal, regular and locally D_K -geodesic. The two instances of additional assumptions for one direction of curvature bounds are decorated by $(*)$ in the figure.

Selected results (1/3)

Theorem [Kunzinger-Sämman 19, Beran-Sämman, 22]

In a strongly causal Lorentzian pre-length space with *timelike curvature bounded below* timelike geodesics do *not branch*.

Theorem [Grant-Kunzinger-Sämman, 19]

A timelike geodesically complete spacetime (or LLS) is *inextendible as a regular LLS*, i.e., any LLS-extension necessarily has unbounded curvature.

Complements class. [Beem-Ehrlich, 80s], C^0 -results [Galloway-Ling-Sbierski, 18] [Minguzzi-Suhr, 19] and relates to *curvature blow up*!

Splitting theorem [Beran-Ohanyan-Rott-Solis, 23]

Let X be a globally hyperbolic LLS with global timelike $K \geq 0$. If X contains a complete timelike line (& some technical conditions) then it splits into a product $\mathbb{R} \times S$ with S a metric length space with $K \geq 0$.

Generalises smooth Lorentzian & synthetic Riemannian results.

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[Grant-Kunzinger-Ohanyan-Schinnerl-S, 25]

- classical Jacobi field estimates
- bound on L of geodesic to 1st conjugate pt.

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$$0 < \text{Sec} \leq K \Rightarrow L_h \geq \pi/\sqrt{K}$$

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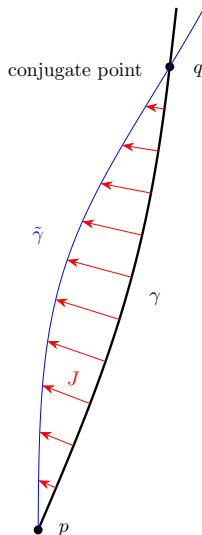
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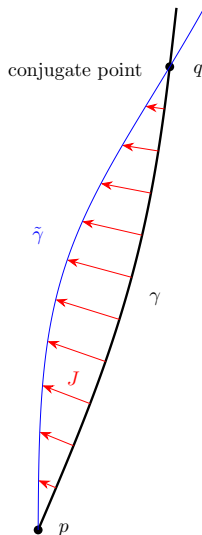
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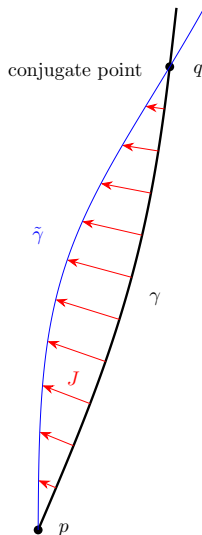
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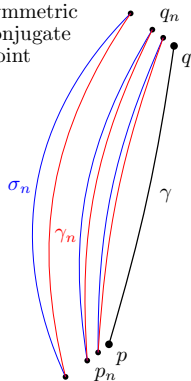
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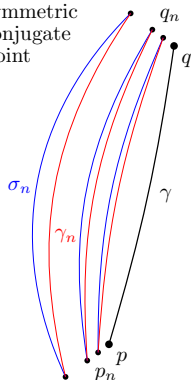
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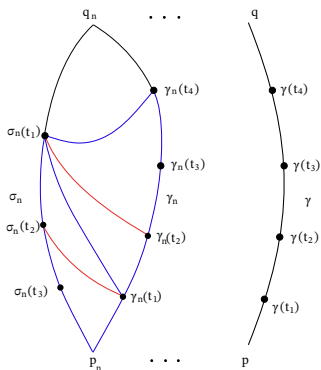
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Closely related to

Synthetic Lorentzian Cartan-Hadamard

[Erös-Gieger, 25]

Let X be a globally hyperbolic regular Lorentzian length space. If X is locally concave and future one-connected, then every pair of timelike related points is connected by a unique timelike geodesic, and these geodesics vary continuously with their endpoints.

Selected results (3/3)

- *Time functions* [Burtscher-García-Heveling, 21]
- *Null distance & LLS* [Kunzinger-S., 22]
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- 2 Lorentzian Geometry & General Relativity
- 3 Smooth Lorentzian Geometry is not good enough
- 4 Interlude: Back to the big picture
- 5 Sectional curvature:
From smooth Riemannian to synthetic Lorentzian
- 6 Ricci curvature:
From smooth Riemannian to synthetic Lorentzian
- 7 Summary & Outlook

Ricci bounds via optimal transport: the basic idea

- *Optimal Transport*: [Monge, Kantorovich]
move matter (distribution μ_1) in the cheapest/optimal way (to μ_2)

- *Minimize*

$$\int_{X \times Y} c(x, y) \, d\pi(x, y) \quad (c \dots \text{cost function})$$

over transport plans $\pi \in \mathcal{P}(X \times Y)$

w given marginals $(\text{pr}_X)_\# \pi = \mu_1$, $(\text{pr}_Y)_\# \pi = \mu_2$

- What is *optimal* depends on cost,
but also on *distances* & *geometry*!
- Turn this on its head:
define *curvature* by requiring that
OT behaves as in model spaces
 - Riemannian case: cost $c = d$
 - Lorentzian case: cost $c = \tau$

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$$\int_{X \times Y} c(x, y) d\pi(x, y) \quad (c \dots \text{cost function})$$

over transport plans $\pi \in \mathcal{P}(X \times Y)$

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- What is *optimal* depends on cost,
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- Turn this on its head:
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Ricci bounds via optimal transport: the basic idea

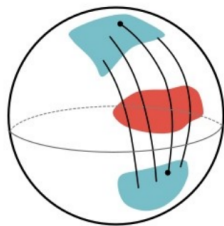
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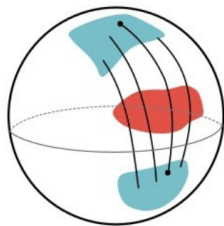
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Ricci Bounds via Optimal Transport: Riemannian case

Thm. (Displacement convexity)

[Lott-Villani, Sturm 06-09]

(M, h) complete Riemannian manifold

$\text{Ric} \geq 0 \iff (M, d_h, \text{vol}_h)$ is an $\text{CD}(0, \infty)$ -space

Definitions. On a metric measure space $(X, d, \mathfrak{m})$ define

- *Wasserstein distance:* $W_2(\mu_1, \mu_2) = \left(\inf_{\pi \in \Pi} \int_{X \times X} d(x, y)^2 d\pi(x, y) \right)^{\frac{1}{2}}$
- *Wasserstein geodesic:* continuous curve $(\mu_t)_{0 \leq t \leq 1}$

$$W_2(\mu_s, \mu_t) = |t - s| \cdot W_2(\mu_1, \mu_2)$$

- *Entropy functional:* $E(\mu|\mathfrak{m}) = \int_X \rho \log \rho d\mathfrak{m}$ for $\mu = \rho\mathfrak{m}$
- $\text{CD}(0, \infty)$ -space: E convex along Wasserstein geodesics, i.e.,

$$E(\mu_t|\mathfrak{m}) \leq (1 - t)E(\mu_1|\mathfrak{m}) + tE(\mu_2|\mathfrak{m})$$

Again turn this into definition of synthetic curvature bounds... $\text{CD}(K, N)$

$\leadsto (R)\text{CD-spaces}$: stability u. measured GH-conv. [Ambrosio-Gigli-Savare, 14]

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Ricci Bounds via Optimal Transport: Lorentzian case

Thm. (Lor. displacement convexity) [McCann, Mondino-Suhr, 20]

(M, g) globally hyperbolic n -dim. spacetime

$\text{Ric}(X, X) \geq 0$ for X timelike $\iff (M, \tau, \text{vol}_g)$ is TCD(0, n)-space

Definitions. Measured Lorentzian pre-length space $(X, \mathfrak{m}, \ll, \leq, \tau)$

- OT & causality [Eckstein-Miller, 17]: $\pi \in \Pi_{\leq}$ (concentrated on $\ll$)
- p -Lorentz Wasserstein distance: $(0 < p \leq 1)$ (for reverse Δ -ineq.)

$$l_p(\mu_1, \mu_2) = \left(\sup_{\pi \in \Pi_{\leq}} \int_{X \times X} \tau(x, y)^p d\pi(x, y) \right)^{1/p}$$

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$$e''(t) - 1/N e'(t)^2 \geq K \int_{X \times X} \tau(x, y)^2 \pi(dx dx)$$

Turn into definition: measured Lorentzian pre-length space w TCD(K, N)

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Selected results

- *Ricci curvature bounds* via optimal transport on *null hypersurfaces*
[McCann, 24, Ketterer, 24, Cavalletti-Manini-Mondino, 24]
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- 2 Lorentzian Geometry & General Relativity
- 3 Smooth Lorentzian Geometry is not good enough
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- 5 Sectional curvature:
From smooth Riemannian to synthetic Lorentzian
- 6 Ricci curvature:
From smooth Riemannian to synthetic Lorentzian
- 7 Summary & Outlook

Summary

A new Lorentzian geometry

(Measured) Lorentzian Length Spaces $(X, \mathfrak{m} \ll, \leq, \tau)$

- provide a general mathematical setting for
 - ▶ *sectional* curvature and
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ingredients: causal set $(X, \leq)$, partial order

locally finite: $J(x, y) = \{z : x \leq z \leq y\}$ finite

CS hypothesis: QT of causal sets X ;

(M, g) approximation of X

$$\mathcal{C}(M, \rho) \ni X \longleftrightarrow (M, g)$$

Hauptvermutung of CST

X can be embedded at density ρ_C into distinct spacetimes iff they are “close”.

$(X, \ll, \leq, \tau)$ is a LpLS

Olivier Ricci curvature and stability

[Barton-Borza-Röhrig, ongoing]

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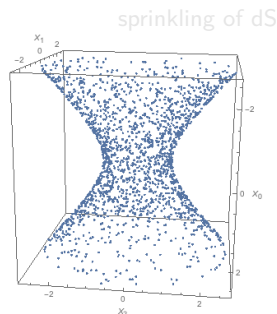
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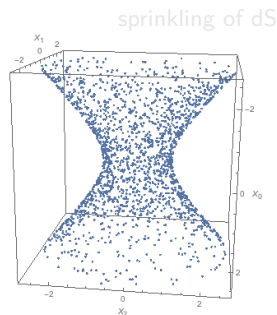
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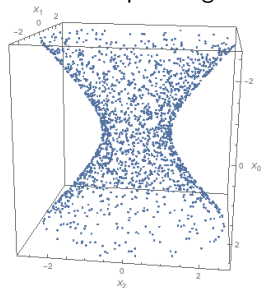
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- chain: $C := (x_i)_{i=1}^n : x_i < x_{i+1}$
- length: $L(C) = n$; $\ll := <$
- $\tau(x, y) := \sup\{L(C) : C \text{ chain from } x \text{ to } y\}$

sprinkling of dS



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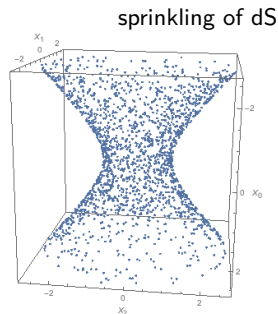
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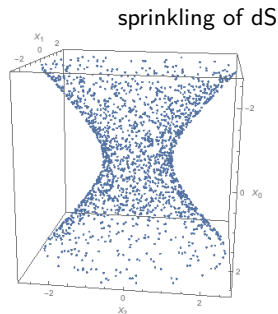
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Meet the EF-Geometry Team



Thank you for your attention

Picture credits

- slide 3, Cartoon-Einstein: <https://wallpapers.com/albert-einstein-png> (freely downloadable)
- slide 3, Spacetime-curvature: Einstein relatively easy, <https://einsteinrelativelyeasy.com/>
- slide 4, lightcone : arXiv:1607.04202[gr-qc], Le Tiec, Alexandre and Novak, Jerome, Theory of Gravitational Waves. arXiv:1607.04202, <https://doi.org/10.48550/arXiv.1607.04202>
- slide 4, cone-trajectory: Stanford Encyclopedia of Philosophy <https://plato.stanford.edu/entries/spacetime-singularities/lightcone.html>
- slide 11, spacetime bubble: Chruściel, Piotr T., James, D. E., On Lorentzian causality with continuous metrics, arXiv:1111.0400 [gr-qc], <https://doi.org/10.48550/arXiv.1111.0400>
- slide 32, spacetime-sprinkling: Sumati Surya, The causal set approach to quantum gravity. Living Reviews in Relativity (2019) 22:5 <https://doi.org/10.1007/s41114-019-0023->
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